

# Numerical values for Nuclear and Quantum Physics

Final version

$\sigma.\beta.$

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## 1 Fundamental constants

$$\begin{aligned}c &= 2.998 \cdot 10^8 \text{ m/s} \\h &= 6.626 \cdot 10^{-34} \text{ Js} \\\hbar &= 1.055 \cdot 10^{-34} \text{ Js} = 6.582 \cdot 10^{-22} \text{ MeV s} \\\hbar c &= 197.327 \text{ MeV fm} \\m_p &= 938.272 \text{ MeV} \\m_n &= 939.565 \text{ MeV} \\\Delta m &= m_n - m_p = 1.293 \text{ MeV} \\m_e &= 0.511 \text{ MeV}\end{aligned}$$

## 2 Atomic constants

$$\begin{aligned}\alpha &= \frac{e^2}{4\pi\epsilon_0\hbar c} = 7.297 \cdot 10^{-3} \approx 1/137 \\\frac{e^2}{4\pi\epsilon_0} &= 1.44 \text{ MeV fm} \\r_e &= \frac{1}{4\pi\epsilon_0} \frac{e^2}{m_e c^2} = 2.818 \text{ fm} \\r_{\text{Bohr}} &= 4\pi\epsilon_0 \frac{\hbar^2}{m_e e^2} = \frac{r_e}{\alpha^2} = 0.529 \text{ \AA} \\Ry &= \frac{1}{(4\pi\epsilon_0)^2} \frac{m_e e^4}{2\hbar^2} = m_e c^2 \alpha^2 / 2 = 13.606 \text{ eV} \\\mu_{\text{Bohr}} &= \frac{e\hbar}{2m_e} = 5.788 \cdot 10^{-11} \text{ MeV/T} \\\mu_N &= \mu_{\text{nuclear}} = \frac{e\hbar}{2m_p} = 3.152 \cdot 10^{-14} \text{ MeV/T}\end{aligned}$$

## 3 Liquid drop model of the nucleus

Binding energy

$$B = a_r A - a_s A^{2/3} - a_c Z^2 / A^{1/3} - a_{\text{sym}} |N - Z|^2 / A + \delta(A)$$

with

$$\begin{aligned}a_r &= 15.69 \text{ MeV} \\a_s &= 16.56 \text{ MeV} \\a_c &= 0.717 \text{ MeV} \\a_{sym} &= 28 \div 30 \text{ MeV} \\a_p &\approx 34 \text{ MeV}\end{aligned}$$

where the pairing effects is given by:

$$\delta(A) = \begin{cases} a_p A^{-3/4} & \text{pari-pari} \\ 0 & \text{disp-pari} \\ -a_p A^{-3/4} & \text{disp-disp} \end{cases}$$

Please note that there are different definitions for the power of  $A$  in  $\delta(A)$  in books, this should be the one accepted in our classes.

## 4 Deutone aka Deuteron

Nucleus of deuterium

$$\begin{aligned}J^P &= 1^+ \\E = -B &= -2.225 \text{ MeV} \\R_d^{rms} &= 2.1 \text{ fm} \\Q_d &= 0.00282 \text{ barn} \\Q_d / (4\pi R_d^2) &\approx 1/196 \sim 0.5\%\end{aligned}$$

It is the mixing of  $^3S_1$  and  $^3D_1$  (the latter accounting for roughly 3.9%).

## 5 Magnetic dipoles

Orbital:

$$\mu_l = g_l \frac{e\hbar}{2m} \frac{l}{\hbar}$$

Spin:

$$\begin{aligned}\mu_s &= g_s \frac{e\hbar}{2m} \frac{s}{\hbar} \\ \mu_{tot} &= \mu_l + \mu_s\end{aligned}$$

In the case of neutron and proton:

$$\mu_N^p = \frac{e\hbar}{2m_p} \approx \frac{e\hbar}{2m_n} = \mu_N^n$$

Orbital gyromagnetic factors:

$$\begin{aligned}g_l^n &= 0 \rightarrow \mu_l^n = 0 \\ g_l^p &= 1 \rightarrow \mu_l^p = 1 \cdot \mu_N \frac{l}{\hbar} \\ g_l^e &= 1 \rightarrow \mu_l^e = 1 \cdot \mu_{Bohr} \frac{l}{\hbar}\end{aligned}$$

Spin g-factors (nucleons and  $e^-$  have  $s = \hbar/2$ )

$$\begin{aligned}g_s^n &= -3.826 \rightarrow \mu_s^n = (g_s^n/2)\mu_N = 2.79\mu_N \\ g_s^p &= 5.586 \rightarrow \mu_s^p = (g_s^p/2)\mu_N = -1.91\mu_N \\ g_s^e &= -2.002 \rightarrow \mu_s^e = (g_s^e/2)\mu_{Bohr} = -\mu_{Bohr}\end{aligned}$$

## 6 Scattering between nucleons

Scattering length

$$a = -\lim_{k \rightarrow 0} \frac{\tan \delta_0}{k}$$

### 6.1 Proton-neutron

$$a_{sing} = (-23.55 \pm 0.12) \text{ fm} < 0 \rightarrow \text{no bound state}$$

$$a_{trip} = (5.35 \pm 0.06) \text{ fm} > 0 \rightarrow \text{weak bound state}$$

Fitting experimental data with Bethe's formula for the effective radius  $r_0$ :

$$k \cotan \delta_0 = -\frac{1}{a} + \frac{1}{2}r_0k^2 + o(k^4)$$

Singlet:

$$a_s^* = (-23.715 \pm 0.015) \text{ fm}$$

$$r_0^s = (2.73 \pm 0.03) \text{ fm}$$

Triplet:

$$a_t^* = (5.423 \pm 0.05) \text{ fm}$$

$$r_0^t = (1.748 \pm 0.006) \text{ fm}$$

### 6.2 Proton-proton

Scattering due to both the electric and nuclear interactions:

$$\text{Nuclear+Electric: } \begin{cases} a^{pp'} = (-7.82 \pm 0.01) \text{ fm} \\ r_o^{pp'} = (2.79 \pm 0.02) \text{ fm} \end{cases}$$

If the electric interaction is removed:

$$\text{Nuclear only: } \begin{cases} a^{pp} = (-17.1 \pm 0.2) \text{ fm} \\ r_o^{pp} = (2.84 \pm 0.03) \text{ fm} \end{cases}$$

### 6.3 Neutron-neutron

$$\begin{cases} a^{nn} = (-16.6 \pm 0.5) \text{ fm} \\ r_o^{nn} = (2.66 \pm 0.15) \text{ fm} \end{cases}$$

## 7 Mean life for common particles

### 7.1 Baryons

$$\tau_p \sim 10^{29} \text{ yrs}$$

$$\tau_n = 880 \text{ s} \sim 15 \text{ min}$$

### 7.2 Leptons

$$m_{\mu\pm} = 105.658 \text{ MeV}$$

$$\tau_{\mu\pm} = 2.2 \text{ } \mu\text{s}$$

$$m_{\tau\pm} = 1776 \text{ MeV}$$

$$\tau_{\tau\pm} = 290 \cdot 10^{-15} \text{ s}$$

## 7.3 Mesons

$$m_{\pi^\pm} = 140 \text{ MeV}$$

$$\tau_{\pi^\pm} = 0.26 \text{ ns}$$

$$m_{\pi^0} = 134.9 \text{ MeV}$$

$$\tau_{\pi^0} = 8.5 \cdot 10^{-17} \text{ s}$$

$$m_{K^\pm} = 494 \text{ MeV}$$

## 8 Useful facts on accelerators

### 8.1 Luminosity

Facility	Type	$\mathcal{L}$	Energy
<i>Daphne (Frascati)</i>	$e^+e^-$	$500 \mu b^{-1}s^{-1}$	1 GeV
<i>Tevatron (Stanford)</i>	$p\bar{p}$	$25 \mu b^{-1}s^{-1}$	2 TeV
<i>LHC (Geneva)</i>	$pp$	$10^6 b^{-1}s^{-1}$ in 2010	14 TeV

### 8.2 CERN's acceleration chain

Accelerator	Exit Energy per beam
<i>LINAC 2</i>	50 MeV
<i>PSB</i>	1.4 GeV
<i>PS</i>	25 GeV
<i>SPS</i>	450 GeV
<i>LHC</i>	8 TeV

## 9 Energy loss for a charge moving in matter

In Bethe-Bloch region, the minimum ionization point is in  $\beta\gamma \sim 3$  which gives a stopping power of  $dE/dx \sim -(1.5 \div 2) \text{ MeV/cm}$  for unitary density.

Cerenkov visible radiation emitted by a particle at  $\beta = 1$  in glass-like material ( $n \sim 1.3$ ) is roughly  $1 \text{ keV/cm}$ .

## 10 Radiation vs matter

- Ultra low energy: photoelectric effect usually for  $E_\gamma \sim 10 \div 100 \text{ eV}$ ;
- Low energy: Rayleigh elastic scattering, for photons whose  $\lambda \gg d$  dimension of target particles (in gas or liquid). Typical values for  $d \sim 1 \text{ \AA}$  are  $E_\gamma \sim 1 \div 10 \text{ eV}$  (visible light) or higher;
- Mid-low energy: Thomson elastic scattering ( $h\nu \ll mc^2$ ) on a target mass  $m$ . If using electrons,  $E_\gamma \ll 500 \text{ keV}$ , i.e. Thomson should dominate till  $E_\gamma \sim 50 \text{ keV}$ ;
- High energy: Compton anelastic scattering on a weakly bound electron, the photon having  $\sim 5 \cdot 10^2 \text{ keV}$ ;
- Ultra high energy:  $e^+e^-$  pair production, for  $\gamma$  with  $E_\gamma > 2m_e \sim 1 \text{ MeV}$  on static nuclei or  $E_\gamma > 4m_e$  on electrons in rest.

## 11 Electrons in matter: Moeller scattering

All the machinery for energy loss in matter depends on the hypothesis that the projectile has  $m \ll M$  of the target. When considering a target mass comparable to the electron's one, Bethe-Bloch and Bremsstrahlung do not work anymore. There still exists a critical energy  $E_c$  such that for  $e^-$  above this value, radiation loss prevail on ionization. For a relativistic electron, we have in different materials:

<i>Material</i>	$\beta\gamma = \gamma$	$E_c$
<i>Pb</i>	20	10 MeV
<i>Air</i>	200	100 MeV

## 12 Radiation length

When electrons lose energy by Bremsstrahlung or pair production, there exists a characteristic length over which the  $e^-$  reduces its energy to  $(1/e)E_0$ , corresponding to 7/9 of the path for creating a couple  $e^-e^+$ .

<i>Material</i>	$X_0$ (cm)	$X'_0$ (g/cm <sup>2</sup> )
<i>Pb</i>	0.56	6.37
<i>Fe</i>	1.76	13.84
<i>Ur</i>	0.32	6.00
<i>Air</i>	30000	36.7

## 13 Strength of interactions

Some constants are still running, but we'll catch'em one day.

$$\begin{aligned}\alpha_{EM} &\sim 1/137 \\ \alpha_{strong} &\sim 1 \\ \alpha_{weak} &\sim 10^{-5} \text{ for } m_p \\ \alpha_{gravity} &\sim 10^{-39}\end{aligned}$$

Most of the values above are taken from the PDG booklet or the notes of our classes of Fisica 3. No responsibility is taken for possible mistakes. Corrections and suggestions are welcome.